

## The Fayans energy-density functional. New constraints from the equations of state.

Friday, 15 July 2022 13:10 (20 minutes)

The Fayans energy-density functional.  
New constraints from the equations of state.

I.N. Borzov 1,2, S.V. Tolokonnikov1,3

1 National Research Centre "Kurchatov Institute", Moscow, Russia

2Bogolubov Laboratory of Theoretical Physics, Joint Institute of Nuclear Research, Dubna, Russia

3 Moscow Institute of Physics and Technology (National Research University), Dolgoprudny, Russia

†E-mail: Borzov\_IN@nrcki.ru, cc: ibor48@mail.ru

The equations of state for infinite, symmetric nuclear matter and pure neutron matter are analyzed in terms of the Fayans energy density functional parameters:  $a_{+,-1,2}$ ,  $h_{+,-1,2}$ . Fitting procedure of the DF3-a functional [1] is redone involving the previously unused parameter  $h_{-2}$ . Additional constraint is implemented from the upper bound of the giant dipole resonance energy in  $^{208}\text{Pb}$ . A quality of the previous global fit of the Fayans EDF has been kept for the nuclear densities, masses of nuclei, single-particle levels and charge radii. Recently the constraints on symmetry energy and its derivative has been obtained in [6] using the data on nuclear masses, results of ab initio calculations with N3LO,  $\Delta R_{np}$  values from PREXP-II, CREX experiments, as well as the latest data from the radii of neutron stars and registration of gravitational waves. The symmetry energy slope at saturation  $L(\rho_0)$  calculated for different  $h_{-2}$  with the relativistic corrections taken into account (Fig.1) is compared with the error margins derived from the set of restrictions [6]. As it can be seen, for DF3-a, the EOS is softer than the ones obtained from the FANDF0 functional [2], as well as from APR [3], AFDMC [4], N2LO(D2,E1) and N2LO(D2,E $\tau$ ) [5] (Fig.2).

Supported by the grant of Russian Scientific Foundation (RSF 21-12-00061).

Fig.1. The  $L(\rho)$  for symmetric nuclear matter. Calculation with the DF3-a functional for various value of parameter  $h_{-2}$ .

Fig. 2. Energy per nucleon for a symmetrical nuclear matter (SNM) as a function of density. our calculation with the FaNDF0[1], DF3-a[2] as well as for APR [3], AFDMC [4], N2LO[5] functionals.

1. S.V. Tolokonnikov, E.E. Saperstein, Phys. At. Nucl. 74, 1277 (2011).
2. S.A. Fayans, JETP Lett. 68, 169 (1998).
3. A. Akmal, V. R. Pandharipande, and D. G. Ravenhall, Phys. Rev. C 58, 1804 (1998).
4. S. Gandolfi, A. Yu. Illarionov, K. E. Schmidt, F. Pederiva, and S. Fantoni, Phys. Rev. C 79, 054005 (2009).
5. D. Lonardonì, I. Tews, S. Gandolfi, and J. Carlson, arXiv:1912.09411 [nucl-th] (2019).
6. J. Lattimer in "Nuclear Matter Symmetry Energy From Experiment, Theory and Observation", S@INT seminar, Seattle, 2021.

### The speaker is a student or young scientist

No

### Section

1. Nuclear structure: theory and experiment

**Primary author:** BORZOV, Ivan (NIC Kurchatov Institute)

**Presenter:** BORZOV, Ivan (NIC Kurchatov Institute)

**Session Classification:** Nuclear structure: theory and experiment