

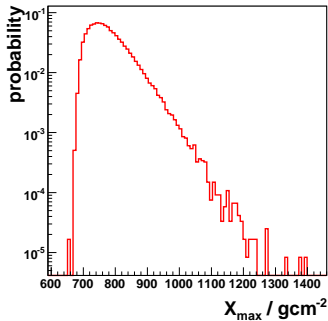
$\sigma(\rho + \text{air})$ with POEMMA



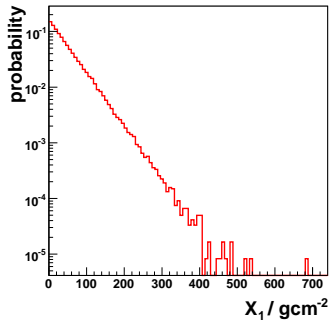
M.H. Reno (U. Iowa), R. Ulrich (KIT), M. Unger (KIT)

Reminder: $\sigma(p + \text{air})$ from X_{\max}

X_{\max} distribution



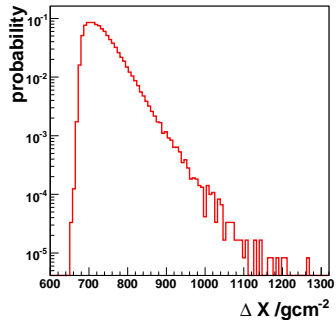
first interaction



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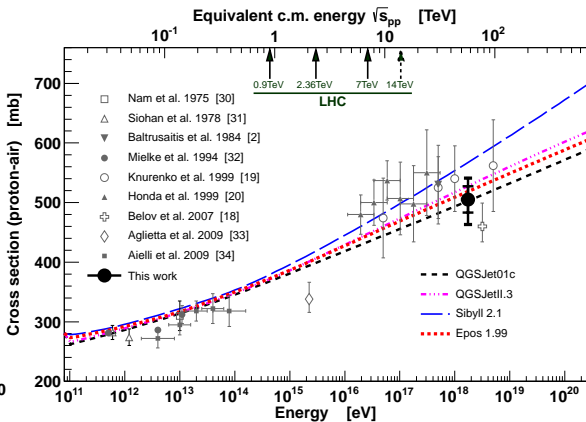
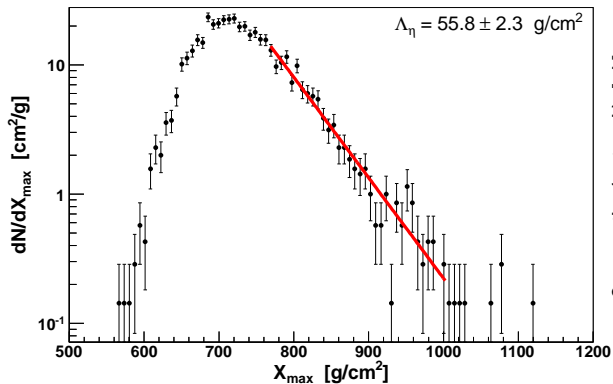


shower development



simulated proton-induced air showers

Reminder: $\sigma(p + \text{air})$ from X_{max}



Fly's Eye Collaboration, PRL 52 (1984) 1380

HiRes Collaboration, Nucl.Phys.Proc.Suppl. 151 (2006) 197

Pierre Auger Collaboration, PRL 109 (2012) 062002

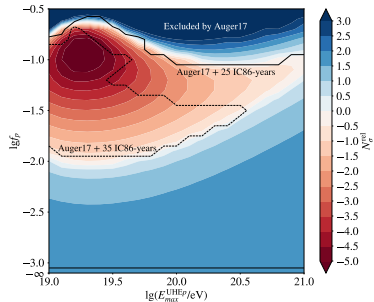
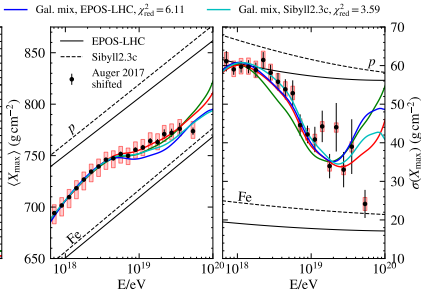
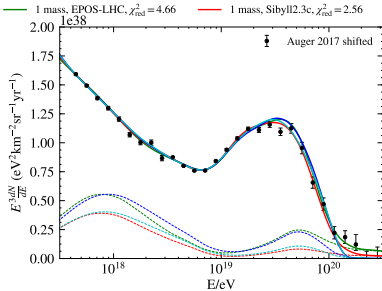
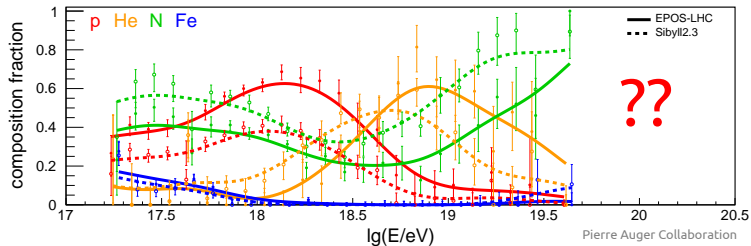
Telescope Array Collaboration, PRD 92 (2015) 32007

Telescope Array Collaboration, arXiv:2006.05012

measured X_{max} distribution $10^{18} \text{ eV} < E < 10^{18.5} \text{ eV}$ (Auger 2012)

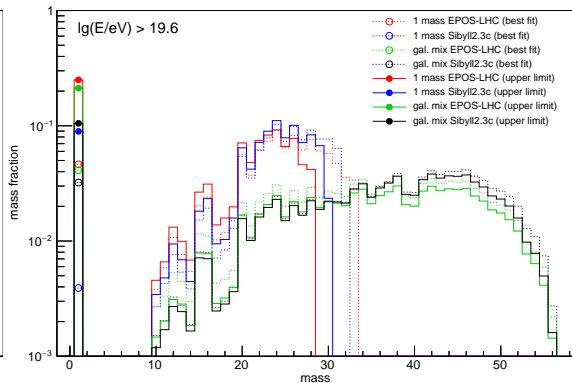
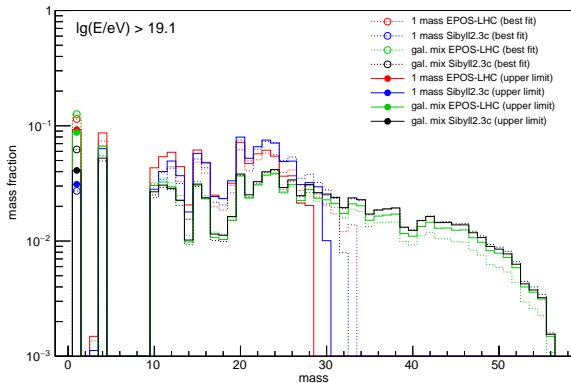
$$\sigma_{p-\text{air}} = \frac{\langle m_{\text{air}} \rangle}{k \Lambda_{\eta}}$$

Are there UHE protons?



UHE Composition Scenarios

- CMB acts as natural He filter (major background for σ_{p-air})
- UHE proton fractions $\lesssim 25\%$ not excluded ($E > 40$ EeV)



POEMMA UHECR Group, Phys.Rev.D 101 (2020) 2, 023012, fractions from M. Muzio et al, Phys.Rev.D 100 (2019) 103008

POEMMA Sensitivity

- $N = 1400$ events $E > 40$ EeV (Auger energy scale)
- relative statistical uncertainty of Λ_η :

$$\sigma_\Lambda/\Lambda = (\eta N)^{-\frac{1}{2}}$$

→ relative statistical uncertainty of $\sigma_{p\text{-air}}$

- two choices of η depending on UHE composition:
 - if $p:N = 1:9 \rightarrow \eta = 0.02$
 - if $p:Si = 1:3 \rightarrow \eta = 0.13$

